

Einstein model of the density of states

Consider N oscillators with the same frequency ω_0 , 1D case.

$$D(\omega) = N \delta(\omega - \omega_0)$$

δ - is the Dirac δ -function centered at ω_0



The thermal energy of the system

$$\begin{aligned} U &= \int d\omega D(\omega) \langle n(\omega) \rangle \hbar\omega \\ &= \int d\omega N \delta(\omega - \omega_0) \frac{\hbar\omega}{e^{\hbar\omega/k_B T} - 1} \end{aligned}$$

$$U = N \hbar\omega_0 \frac{1}{e^{\hbar\omega_0/k_B T} - 1}$$

The heat capacity $C_V = \frac{\partial U}{\partial T}$

$$C_V = \left(\frac{\partial U}{\partial T} \right) = N k_B \left(\frac{\hbar\omega_0}{k_B T} \right)^2 \frac{e^{\hbar\omega_0/k_B T}}{(e^{\hbar\omega_0/k_B T} - 1)^2}$$

obtained in 1907 (Einstein); contribution from N identical oscillators to the heat capacity

$$\text{In 3D: } C_V = \left(\frac{\partial U}{\partial T} \right) = 3N k_B \left(\frac{\hbar\omega_0}{k_B T} \right)^2 \frac{e^{\hbar\omega_0/k_B T}}{(e^{\hbar\omega_0/k_B T} - 1)^2}$$

Check that $T \rightarrow \infty$,

$$C_V = 3Nk_B$$

Classical value
of the heat
capacity

Dulong-Petit
value

$$C_V = 3Nk_B \left(\frac{\theta_E}{T} \right)^2 \frac{e^{\theta_E/T}}{(e^{\theta_E/T} - 1)^2}$$

$$\theta_E = \frac{\hbar \omega_0}{k_B}$$

General result for $D(\omega)$

We want to find a general expression for $D(\omega)$

of states per unit frequency range

given that we have a dispersion relation $\omega(\vec{k})$

→ The number of allowed \vec{k} -values for which the phonon frequency lies between ω , $\omega + d\omega$ is given by

$$D(\omega) d\omega = \frac{V}{(2\pi)^3} \int_{\text{shell}} d^3k$$

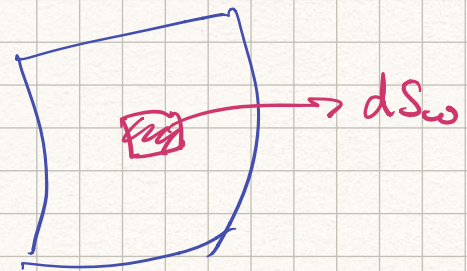
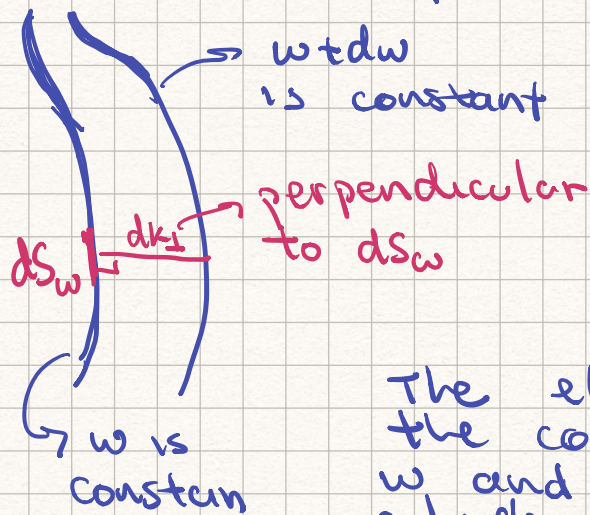
The integration is done over the volume of the shell in \vec{k} -space bounded by two surfaces on which the frequency of phonons is constant

→ one has frequency ω

→ other has frequency $\omega + d\omega$

The problem is to evaluate the volume of the shell.

Let dS_ω denote an element of area on the surface in \vec{k} -space at which ω is constant.



The element of volume between the constant frequency surfaces ω and $\omega + d\omega$ is a right cylinder of base dS_ω and height dk_\perp .

$$\int_{\text{shell}} d\vec{k} = \int dS_\omega dk_\perp$$

dk_\perp is the perpendicular distance between the two surfaces, and it may vary from one point to another.

The gradient of ω , given $\vec{\nabla}_k \omega$, is also \perp to the surface where ω is constant.

$$\vec{\nabla}_k = (\partial_{k_x}, \partial_{k_y}, \partial_{k_z})$$

$|\vec{\nabla}_k \omega| dk_\perp = d\omega$ is the difference in frequency between the two surfaces corrected by dk_\perp .

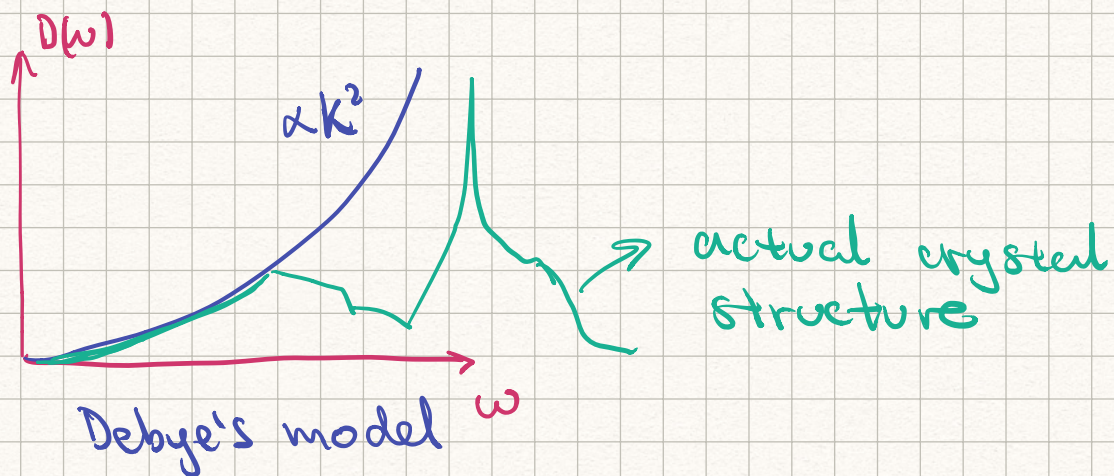
$$dS_\omega dk_\perp = dS_\omega \frac{d\omega}{|\vec{\nabla}_k \omega|} = dS_\omega \frac{d\omega}{|\vec{v}_g|}$$

$|\vec{v}_g| = |\vec{\nabla}_k \omega|$ is the magnitude of the group velocity.

of phonons

$$D(\omega) d\omega = \frac{V}{(2\pi)^3} \int_{\text{shell}} d^3k$$
$$= \frac{V}{(2\pi)^3} \int \frac{ds_\omega}{v_g} d\omega$$

$$D(\omega) = \frac{V}{(2\pi)^3} \int \frac{ds_\omega}{v_g}$$



Free electron Fermi Gas

Energy levels: 1D

An electron confined in a length L in 1D.
The Schrödinger Eq.

$$H\psi_n = E_n\psi_n$$

The Hamiltonian $H = \frac{p^2}{2m}$ $p = -i\hbar\partial_x$

$$\frac{p^2}{2m} = \frac{(-i\hbar\partial_x)^2}{2m} = -\frac{\hbar^2\partial_x^2}{2m}$$

$$-\frac{\hbar^2}{2m} \frac{\partial^2\psi_n}{\partial x^2} = E_n\psi_n$$

We know that the solution to this equation

$$\psi_n(x) = A_n \sin(kx) + B_n \cos(kx)$$

Because the electron is confined to a length L

$$\psi_n(0) = \psi_n(x=L) = 0$$

$$\begin{aligned} \psi_n(x) &= A_n \sin(kx) & kL &= n\pi & n \in \mathbb{Z} \\ &= A_n \sin\left(\frac{n\pi x}{L}\right) & k &= \frac{n\pi}{L} \end{aligned}$$

Replacing in S.E. : $E_n = \frac{\hbar^2 k^2}{2m} = \frac{\hbar^2}{2m} \left(\frac{n\pi}{L}\right)^2$

We want to accommodate N electrons

- Pauli exclusion principle

→ No two electrons can have all their quantum numbers equal.

- The Q.N. that play a role are

n : principal quantum number : Associated to E_n

m_s : magnetic q. number : Associated to the spin

Electrons can have $m_s = \pm 1/2$; $\left\{ \uparrow, \downarrow \right\}$ "spin"

Example: we have 8 electrons

n	m_s	Occupation (how many e^-)
1	\uparrow	1
1	\downarrow	1
2	\uparrow	1
2	\downarrow	1
3	\uparrow	1

3	↓	1	n_F : Fermi level
4	↑	1	Topmost filled energy level of the system
4	↓	1	
5	↑	0	→ we start filling at the bottom ($n=1$).
5	↓	0	→ all e^- are accommodated

In this case $n_F = 4$

→ all e^- are accommodated

If N is an even number $\Rightarrow n_F = N/2$: This cond. determines n_F

E_F : Fermi energy

Energy of the topmost filled level in the ground state of the N electron system

$$E_F = \frac{\hbar^2}{2m} \left(\frac{n_F \pi}{L} \right)^2 = \frac{\hbar^2}{2m} \left(\frac{N\pi}{2L} \right)^2$$

Effect of temperature

→ The ground state is defined at $T=0$ when we increase the temperature

↳ e^- gain kinetic energy

The probability of having an energy level ϵ occupied is given by the Fermi-Dirac distribution

$$f(\epsilon) = \frac{1}{\exp[(\epsilon - \mu)/k_B T] + 1}$$

μ : chemical potential = energy that we need to add/remove a particle
 μ depends on T , and when $T=0$, $\mu = E_F$.

The high energy tail of the distribution for which $\epsilon - \mu \gg k_B T$, is dominated by

$$f(\epsilon) \approx e^{-(\epsilon - \mu)/k_B T}$$

which is the Maxwell-Boltzmann distribution, describing classical particles (no spin).